

Monte Carlo method projective estimators for
angular and temporal characteristics evaluation of
polarized radiation

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Let us consider the following stationary integro-differential radiative transfer equation with polarization:

$$\omega \nabla \Phi(r, \omega) + \sigma \Phi(r, \omega) = \int_{\Omega} \sigma_s P(\omega', \omega) \Phi(r, \omega') d\omega' + \mathbf{f}_0(r, \omega),$$

or in the operator form

$$\mathbf{L}\Phi + \sigma\Phi = \mathbf{S}\Phi + \mathbf{f}_0,$$

where $\Phi = (\Phi_1, \Phi_2, \Phi_3, \Phi_4)^T$ – is a vector function describing the flux density of particles (of “vector photons”), or, in other words, the radiation intensity; Ω – is the space of unit direction vectors, $\omega \in \Omega$, $r \in D \subset \mathbb{R}^3$; $P(\omega', \omega)$ is the matrix scattering function, $\sigma = \sigma(r)$ is the total cross section, $\sigma = \sigma_s + \sigma_c$, σ_c is the absorption cross section, σ_s is the scattering cross section; $\mathbf{f}_0 = (f_0^{(1)}, f_0^{(2)}, f_0^{(3)}, f_0^{(4)})^T$ is the vector distribution density of particles.

The matrix $P(\omega', \omega)$ is defined by the relation

$$P(\omega', \omega) = L(\pi - i_2)R(\omega', \omega)L(-i_1),$$

where $L(\cdot)$ is a special rotation matrix

$$L(i) = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \cos 2i & \sin 2i & 0 \\ 0 & -\sin 2i & \cos 2i & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix},$$

$R(\cdot)$ is the scattering matrix; i_1 is the angle between the plane ω', s and the scattering plane ω, ω' ; i_2 is the angle between the scattering plane ω, ω' and the plane ω, s ; $s = (0, 0, 1)$.

In the case of an isotropic medium, the scattering matrix is of the form:

$$R(\mu, r) = \frac{1}{2\pi} \begin{pmatrix} r_{11} & r_{12} & 0 & 0 \\ r_{21} & r_{22} & 0 & 0 \\ 0 & 0 & r_{33} & r_{34} \\ 0 & 0 & -r_{43} & r_{44} \end{pmatrix},$$

where $\mu = (\omega, \omega')$ is cosine of the scattering angle, r_{11} is the scattering function (or indicatrix), $\int_{-1}^1 r_{11}(\mu) d\mu = 1$. If the scattering particles are homogeneous spheres, then $r_{11} = r_{22}$, $r_{12} = r_{21}$, $r_{33} = r_{44}$, $r_{34} = r_{43}$.

In the case of molecular (Rayleigh) scattering, the matrix R is defined as follows:

$$R(\mu) = \frac{1}{2\pi} \begin{pmatrix} 3(1 + \mu^2)/8 & -3(1 - \mu^2)/8 & 0 & 0 \\ -3(1 - \mu^2)/8 & 3(1 + \mu^2)/8 & 0 & 0 \\ 0 & 0 & 3\mu/4 & 0 \\ 0 & 0 & 0 & 3\mu/4 \end{pmatrix}.$$

Angular characteristics

Consider the radiation transport through a flat layer $0 < z < H$ of scattering and capturing media from the radiation source located on the border of $z = 0$ and directed in some direction ω_0 .

Vector-function $\Phi_s(r, \omega)$ of the angular distribution of the scattered radiation flux at r point on the surface $z = h$, $0 \leq h \leq H$ is

$$\Phi_s(r, \mu, \varphi) = |\mu| \Phi(r, \mu, \varphi),$$

where $\mu = (\omega, s)$, φ is the azimuthal angle.

Total radiation flux density (surface illumination) equal to

$$P_h = \int_{\Omega} \Phi^{(1)}(r, \omega) \mu d\omega = \int_0^{2\pi} \int_0^1 \Phi^{(1)}(r, \mu, \varphi) \mu d\mu d\varphi.$$

In the case of isotropic radiation, the total flux density is $P_h^{Lambert} = \pi \Phi^{(1)}(r)$ and the corresponding angular distribution of the radiation flux on a surface with a normalized density

$$\Phi_s^{(1)Lambert}(r, \mu, \varphi) / P_h^{Lambert} \equiv \mu / \pi, \quad \mu \in (0, 1), \quad \varphi \in (0, 2\pi)$$

is well-known Lambert distribution.

Study

- angular distribution of backscattered and transmitted radiation intensity

$$\Phi_s^{(1)}(x, y)/P_H, \quad x \in (0, 1), y \in (0, \pi),$$

- degree of polarization

$$p(r, \omega) = \frac{\sqrt{\Phi^{(2)}(r, \omega)^2 + \Phi^{(3)}(r, \omega)^2 + \Phi^{(4)}(r, \omega)^2}}{\Phi^{(1)}(r, \omega)}.$$

1: $\{\psi_i(x) \cdot \phi_j(y)\}$, $i, j = 0, \dots, \infty$, such that

$$\int_0^{2\pi} \int_0^1 x \psi_i(x) \phi_k(y) \psi_j(x) \phi_l(y) dx dy = \begin{cases} 1, & i = j, k = l \\ 0, & \text{else.} \end{cases}$$

$$\int_0^1 x \psi_i(x) \psi_j(x) dx = \begin{cases} 1, & i = j \\ 0, & i \neq j. \end{cases}$$

$$\int_0^{2\pi} \phi_k(y) \phi_l(y) dy = \begin{cases} 1, & k = l \\ 0, & k \neq l. \end{cases}$$

2: $\{H_{i,j}(x, y), i = 0, \dots, \infty, j = -i, \dots, i\}$ such that

$$\int_0^{2\pi} \int_0^1 x H_{i,k}(x, y) H_{j,l}(x, y) dx dy = \begin{cases} 1, & i = j, k = l \\ 0, & \text{else.} \end{cases}$$

- N. V. Tracheva, S. A. Ukhinov (2017) *Numerical statistical study of the angular distribution of the polarized radiation scattered by medium*// Russian Journal of Numerical Analysis and Mathematical Modelling. Vol. 32(2), 135-146.
- Tracheva, N. V. (2017) *The use of two-dimensional projective Monte Carlo estimators for solution of number of problems of theory of radiation transfer*// Proceedings of the Conference of young scientists of Institute of computational mathematics and mathematical geophysics of SB RAN. Novosibirsk, 52–63.

Functions $\psi_i(x)$ were obtained from Jacobi polynomials:

$$\psi_i(x) = \sqrt{2i+2} \sum_{k=0}^i \frac{(-1)^k (2i+1-k)!}{(i-k)! k! (i+1-k)!} x^{i-k}.$$

Functions $\phi_j(x)$ were obtained from Legendre polynomials:

$$\phi_j(y) = \frac{\sqrt{2j+1} (-1)^j}{\sqrt{2\pi}} \sum_{k=0}^j \frac{(j+k)!}{(j-k)! k! k! (2\pi)^k} (-x)^k.$$

- Mahotkin O. A. (1996) *Analysis of radiative transfer between surfaces by hemispherical harmonics*// Journal of Quantitative Spectroscopy and Radiative Transfer. Vol. 56(6), 869-879.

The basis is constructed on the basis of hemispherical functions $\{H_{i,j}(\mu, \varphi)\}$ of a form

$$H_{0,0}(\mu, \varphi) = \frac{1}{\sqrt{\pi}}; \quad H_{i,0}(\mu, \varphi) = \frac{\sqrt{i+1}}{\sqrt{\pi}} \sum_{k=0}^i \frac{(-1)^k (2i+1-k)!}{(i-k)! k! (i+1-k)!} \mu^{i-k}, \quad i = 1, 2, \dots;$$

$$H_{i,j}(\mu, \varphi) = \frac{\sqrt{2(i+1)}}{2^j \sqrt{\pi \cdot \prod_{k=1}^j [(i+1)^2 - k^2]}} \cdot \cos(j\varphi) \times$$

$$\times \sum_{k=0}^{i-j} \frac{i!(i+j+k+1)!(i+j+1)!(\mu-1)^k (1-(2\mu-1)^2)^{j/2}}{k!(l+j+1)!(j+k)!(i-j-k)!(i+1)!}, \quad i = 1, 2, \dots, j = 1, \dots, i;$$

$$H_{i,-j}(\mu, \varphi) = \frac{\sqrt{2(i+1)}}{2^j \sqrt{\pi \cdot \prod_{k=1}^j [(i+1)^2 - k^2]}} \cdot \sin(j\varphi) \times$$

$$\times \sum_{k=0}^{i-j} \frac{i!(i+j+k+1)!(i+j+1)!(\mu-1)^k (1-(2\mu-1)^2)^{j/2}}{k!(l+j+1)!(j+k)!(i-j-k)!(i+1)!}, \quad i = 1, 2, \dots, j = 1, \dots, i.$$

Vector-function of the angular distribution of the scattered radiation flux on the surface $z = h$, $0 \leq h \leq H$:

$$\Phi_s(x, y) = x \sum_{i=0}^{\infty} \sum_{j=-i}^i \mathbf{a}_{i,j} H_{i,j}(x, y),$$

where

$$\mathbf{a}_{i,j} = \int_0^{2\pi} \int_0^1 \Phi_s(x, y) H_{i,j}(x, y) dx dy.$$

$$\xi_{i,j} = \sum_{k=1}^{N_t-1} q_k \mathbf{Q}_{k+1} H_{i,j}(\mu_{k+1}, \varphi_{k+1}) e^{-\frac{\sigma(h-z_k)}{\mu_{k+1}}} \Delta_{s,\mu}(z_k, \mu_{k+1}), \quad E\xi_{i,j} = a_{i,j}.$$

$$\xi_h = \sum_{k=1}^{N_t-1} q_k \mathbf{Q}_{k+1} e^{-\frac{\sigma(h-z_k)}{\mu_{k+1}}} \Delta_{s,\mu}(z_k, \mu_{k+1}), \quad E\xi_h = P_h.$$

constructed on a Markov chain $\{x_k\}$, where $x_k = (z_k, \omega_k)$, z_k – z -coordinate of the collision point, $z_0 = 0$; ω_k – the direction of the particle before the collision event, $\omega_1 = \omega_0$; $\mu_k = (\omega_k, \mathbf{e}_z)$;

$$\Delta_{s,\mu}(z_k, \mu_{k+1}) = \begin{cases} 1, & \text{if } z_k < h \text{ and } \mu_{k+1} > 0 \\ 0, & \text{else,} \end{cases}$$

i.e. $\Delta_{s,\mu}$ – the indicator of the surface $z = h$ intersection by the particle direction after the scattering event in the point $z_k < h$, N_t – random number of the last collision event. Direction of the scattering ω_k in point z_{k-1} simulated according to the element R_{11} of the scattering matrix $R(\omega_{k-1}, \omega_k)$. Vector weight $\mathbf{Q}_k = (Q_k^{(1)}, Q_k^{(2)}, Q_k^{(3)}, Q_k^{(4)})^T$ defines by:

$$\mathbf{Q}_1 = (I_0^{(1)}, I_0^{(2)}, I_0^{(3)}, I_0^{(4)})^T,$$

$$\mathbf{Q}_k = P(\omega_{k-1}, \omega_k) \mathbf{Q}_{k-1} / R_{11}(\omega_{k-1}, \omega_k), \quad k \geq 2.$$

$$\Phi(x, y) \approx \sum_{i=0}^n \sum_{j=-i}^i \mathbf{a}_{i,j} H_{i,j}(x, y) = \Phi_n(x, y) \approx \tilde{\Phi}_n(x, y) = \sum_{i=0}^n \sum_{j=-i}^i \alpha_{i,j} H_{i,j}(x, y),$$

where

$$\alpha_{i,j} = E_N \boldsymbol{\xi}_{i,j} \equiv \frac{1}{N} \sum_{m=0}^N \boldsymbol{\xi}_{i,j}^{(m)},$$

$\boldsymbol{\xi}_{i,j}^{(m)}$ – random value realization $\boldsymbol{\xi}_{i,j}$ on m -th trajectory,

$$E \alpha_{i,j} = E \boldsymbol{\xi}_{i,j} = \mathbf{a}_{i,j},$$

$$E \tilde{\Phi}_n(x, y) = \Phi_n(x, y).$$

Componentwise representation of variation of random vector-function $\tilde{\Phi}_n(x)$

$$D\tilde{\Phi}_n(x, y) = \sum_{i=0}^n \sum_{k=-i}^i \sum_{j=0}^n \sum_{l=-j}^j \text{cov}(\alpha_{i,k}, \alpha_{j,l}) H_{i,k}(x, y) H_{j,l}(x, y) =$$

$$\frac{1}{N} \sum_{i=0}^n \sum_{k=-i}^i \sum_{j=0}^n \sum_{l=-j}^j \text{cov}(\xi_{i,k}, \xi_{j,l}) H_{i,k}(x, y) H_{j,l}(x, y).$$

$$\tilde{p}_n(x, y) = \frac{\sqrt{(\tilde{\Phi}_n^{(2)}(x, y))^2 + (\tilde{\Phi}_n^{(3)}(x, y))^2 + (\tilde{\Phi}_n^{(4)}(x, y))^2}}{\tilde{\Phi}_n^{(1)}(x, y)}.$$

$$\begin{aligned}
D_N \tilde{p}_n(x) &= \frac{1}{g_1^2} \left(\frac{g D_N g_1}{g_1^2} + \frac{g_2^2 D_N g_2 + g_3^2 D_N g_3 + g_4^2 D_N g_4}{g} \right. \\
&\quad \left. - 2 \frac{\text{cov}_N(g_1, g_2) g_2 + \text{cov}_N(g_1, g_3) g_3 + \text{cov}_N(g_1, g_4) g_4}{g_1} + \right. \\
&\quad \left. + 2 \frac{\text{cov}_N(g_2, g_3) g_2 g_3 + \text{cov}_N(g_2, g_4) g_2 g_4 + \text{cov}_N(g_3, g_4) g_3 g_4}{g} \right) + \\
&\quad + o(N^{-1}),
\end{aligned}$$

where $g_m = \tilde{\Phi}_n^{(m)}(x, y)$, $g = g_2^2 + g_3^2 + g_4^2$,

$$\text{cov}_N(g_m, g_t) = \frac{1}{N} \sum_{i=0}^n \sum_{k=-i}^i \sum_{j=0}^n \sum_{l=-j}^j \text{cov}_N(\xi_{i,k}^{(m)}, \xi_{j,l}^{(t)}) H_{i,k}(x, y) H_{j,l}(x, y), \quad m, t = 1, \dots, 4.$$

Temporal characteristics

An exponential asymptotics parameter is the principal eigenvalue λ^* of the homogeneous stationary transfer equation:

$$\mathbf{L}\Phi + (\sigma + \lambda/v)\Phi = \mathbf{S}\Phi,$$

with standard boundary conditions (see Davison B. *Neutron Transport Theory*. Clarendon, Oxford. 1957.).

In the work of Mikhailov G. A., Tracheva N. V., Uhinov S. A. (Monte Carlo study of time asymptotics of the polarized radiation intensity // *Computational Mathematics and Mathematical Physics*. 2007. V. 47, N. 7.) this statement was extended to polarized radiation. In the spatially homogeneous case (i.e., when the entire space is filled with a homogeneous medium), this goal can be achieved rather easily through weighted simulation of radiative transfer. Moreover, it proves that $\lambda^* = \lambda_{\infty}^* = -\sigma_c v$ irrespective of the polarization type. Here, σ_c is the absorption cross section and v is the velocity of the particles.

It is well known that $\lambda^* = \lambda_{\infty}^* = -\sigma_c v$ for a half-space as well.

A number of publications, e.g.

- Romanova L. M. (1965) Limit cases of the free path distribution function of photons exiting a thick light-diffusing layer.// *Izv. Akad. Nauk SSSR, Ser. Fiz. Atmos. Okeana.*, **1**, No. 6, 599 - 606.
- Zege E.P., Katsev I.L. (1973) *Time asymptotical solutions of radiation transfer equation and their application*. Preprint. Inst. Fiz. Akad. Nauk. BSSR, Minsk.

show that for intensity functionals without polarization the following asymptotic relation exists:

$$J(r, \omega, t) \sim C(r, \omega) \cdot t^\alpha \cdot e^{-\sigma_c vt}, \quad t \rightarrow +\infty.$$

We are concerned with obtaining of the similar relation in pre-asymptotical times.

Let us consider a system $\psi_i(t)$ of orthonormal over $(0, \infty)$ with weight $e^{-\sigma_c vt}$ functions, such that

$$\int_0^{\infty} e^{-\sigma_c vt} \psi_i(t) \psi_j(t) dt = \begin{cases} 1, & i = j, \\ 0, & \text{else.} \end{cases}$$

$$\Phi(r, \omega, t) = C(r, \omega) e^{-\sigma_c vt} \sum_{i=0}^{\infty} a_i \psi_i(t),$$

where

$$a_i = C(r, \omega) \int_0^{\infty} \Phi(t) \psi_i(t) dt.$$

$$\xi_i = \sum_{k=1}^{N_t-1} q_k \mathbf{Q}_{k+1} \psi_i(t_{k+1}), \quad E\xi_i = a_i$$

built on Markov chain $\{x_k\}$, $x_k = (r_k, \omega_k)$, r_k – coordinates of the collision point, $r_0 = (0, 0, 0)$; ω_k – is the direction of the particle before the collision, $\omega_1 = \omega_0$; $\mu_k = (\omega_k, \mathbf{e}_z)$, N_t – random number of the last collision event, the scattering direction ω_k at the point r_{k-1} is simulated according to an element R_{11} of scattering matrix $R(\omega_{k-1}, \omega_k)$. Vector weight $\mathbf{Q}_k = (Q_k^{(1)}, Q_k^{(2)}, Q_k^{(3)}, Q_k^{(4)})^T$ is determined by

$$\mathbf{Q}_1 = (I_0^{(1)}, I_0^{(2)}, I_0^{(3)}, I_0^{(4)})^T,$$

$$\mathbf{Q}_k = P(\omega_{k-1}, \omega_k) \mathbf{Q}_{k-1} / R_{11}(\omega_{k-1}, \omega_k), \quad k \geq 2.$$

$$\begin{aligned}\Phi(r, \omega, t) &\approx C(r, \omega) e^{-\sigma c v t} \sum_{i=0}^n a_i \psi_i(t) = \Phi_n(r, \omega, t) \approx \\ &\approx \tilde{\Phi}_n(r, \omega, t) = C(r, \omega) e^{-\sigma c v t} \sum_{i=0}^n \alpha_i \psi_i(t),\end{aligned}$$

where

$$\alpha_i = E_N \xi_i \equiv \frac{1}{N} \sum_{m=0}^N \xi_i^{(m)},$$

$\xi_i^{(m)}$ – random value ξ_i realization on m -th trajectory,

$$E \alpha_i = E \xi_i = a_i,$$

$$E \tilde{\Phi}_n(r, \omega, t) = \Phi_n(r, \omega, t).$$

Variation of random function $\tilde{\Phi}_n(t)$ is equal to

$$\begin{aligned} D\tilde{\Phi}_n(t) &= C^2(r, \omega) e^{-2\sigma_c vt} \sum_{i=0}^n \sum_{j=0}^n \text{cov}(\alpha_i, \alpha_j) \psi_i(t) \psi_j(t) = \\ &= C(r, \omega)^2 e^{-2\sigma_c vt} \frac{1}{N} \sum_{i=0}^n \sum_{j=0}^n \text{cov}(\xi_i, \xi_j) \psi_i(t) \psi_j(t). \end{aligned}$$

As orthonormal basis let us consider a modified Laguerre polynomials $L_n(x)$, orthogonal with weight e^{-x} over $(0, \infty)$. We can obtain the following explicit form for functions $\psi_i(t)$:

$$\psi_i(x) = \sqrt{\sigma_c v} \sum_{k=0}^i \frac{(-1)^k i!}{k! (i-k)! k!} \sigma_c^k v^k t^k.$$

Numerical results

- Deirmendjian, D. (1969) *Electromagnetic Scattering on Spherical Polydispersions*. Amer. Elsevier, New York.

Aerosol scattering.

Coefficient of refraction on particles $n = 1.331 - i1.310^{-4}$ (water), particle size distribution is logarithmically normal with density $f(r) = \frac{1}{r} \exp(-\frac{1}{2\sigma_g^2} \ln^2(\frac{r}{r_g}))$, $r \in (0, 10\text{mkm})$, $r_g = 0.12\text{mkm}$, $\sigma_g = 0.5$, wave-length of radiation equals to 0.65 mkm. Mean cosine of scattering angle - $\mu_0 = 0.7292$.

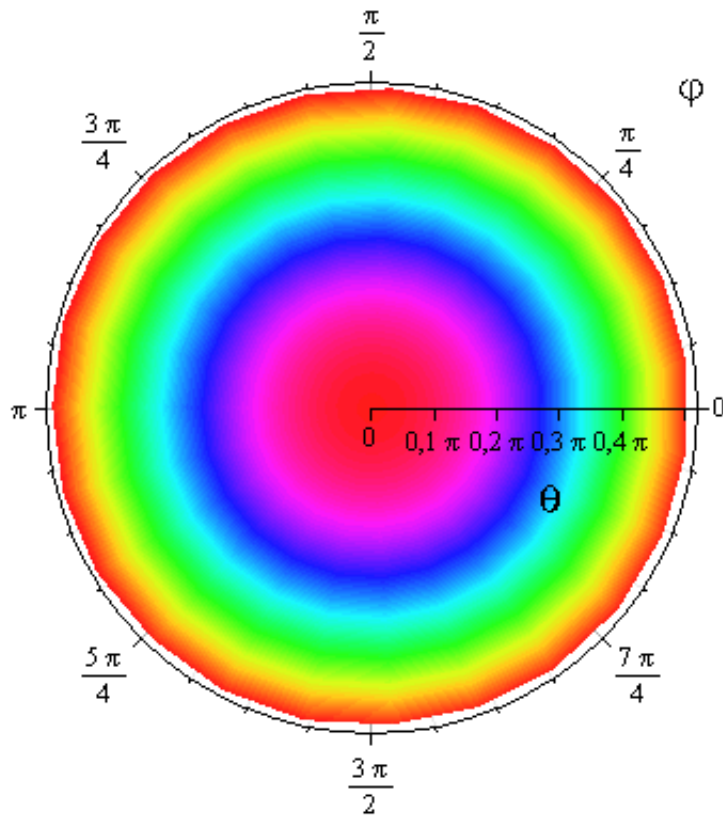
- Marchuk, G. I., Mikhailov, G. A., Nazaraiev, M. A. et al. (1980) *The Monte Carlo Methods in Atmospheric Optics*. Amer. Springer-Verlag, Heidelberg .

Molecular scattering.

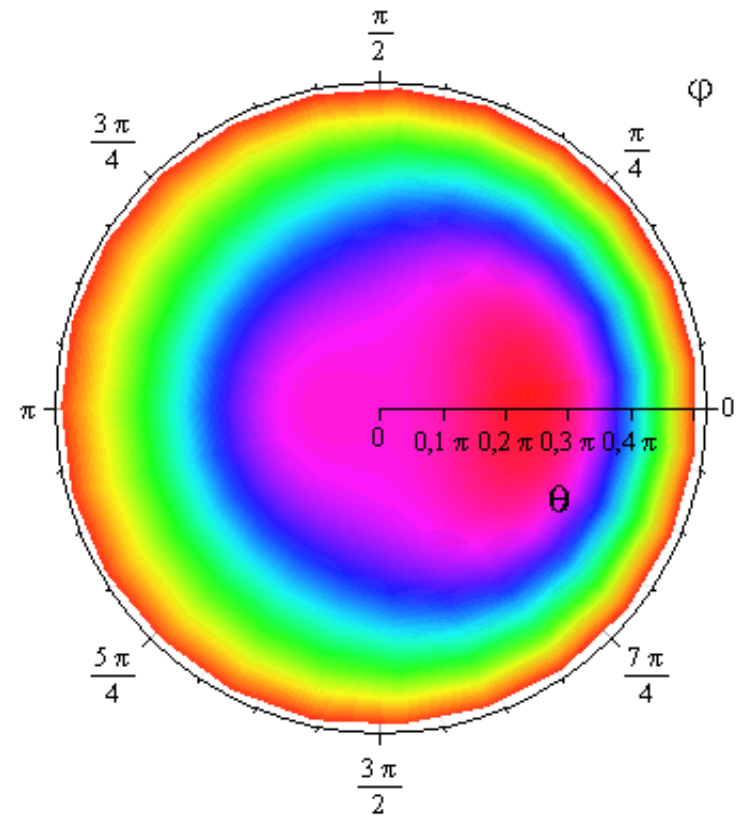
$$R(\mu) = \frac{1}{2\pi} \begin{pmatrix} 3(1 + \mu^2)/8 & -3(1 - \mu^2)/8 & 0 & 0 \\ -3(1 - \mu^2)/8 & 3(1 + \mu^2)/8 & 0 & 0 \\ 0 & 0 & 3\mu/4 & 0 \\ 0 & 0 & 0 & 3\mu/4 \end{pmatrix} .$$

Angular distribution function of the normalized flux of radiation $\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$, backscattered by a layer of optical thickness $H = 5$. Aerosol scattering. $\theta_0 = 45^\circ$, $\varphi_0 = 0$. Hemispherical basis.

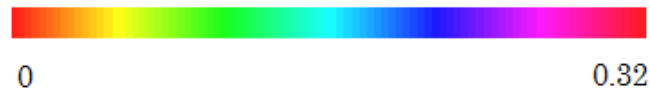
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Lambertian approximation

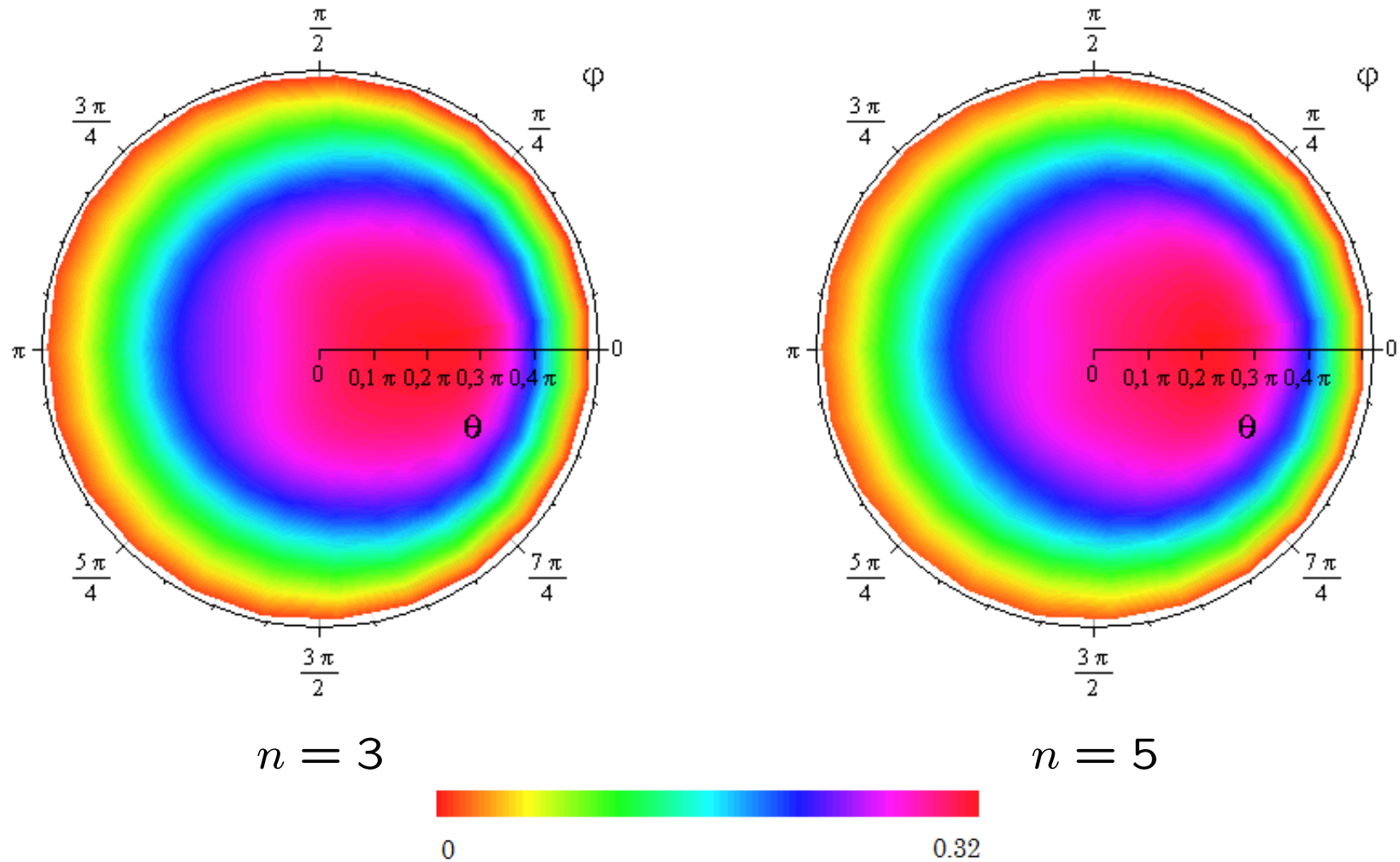


$n = 2$

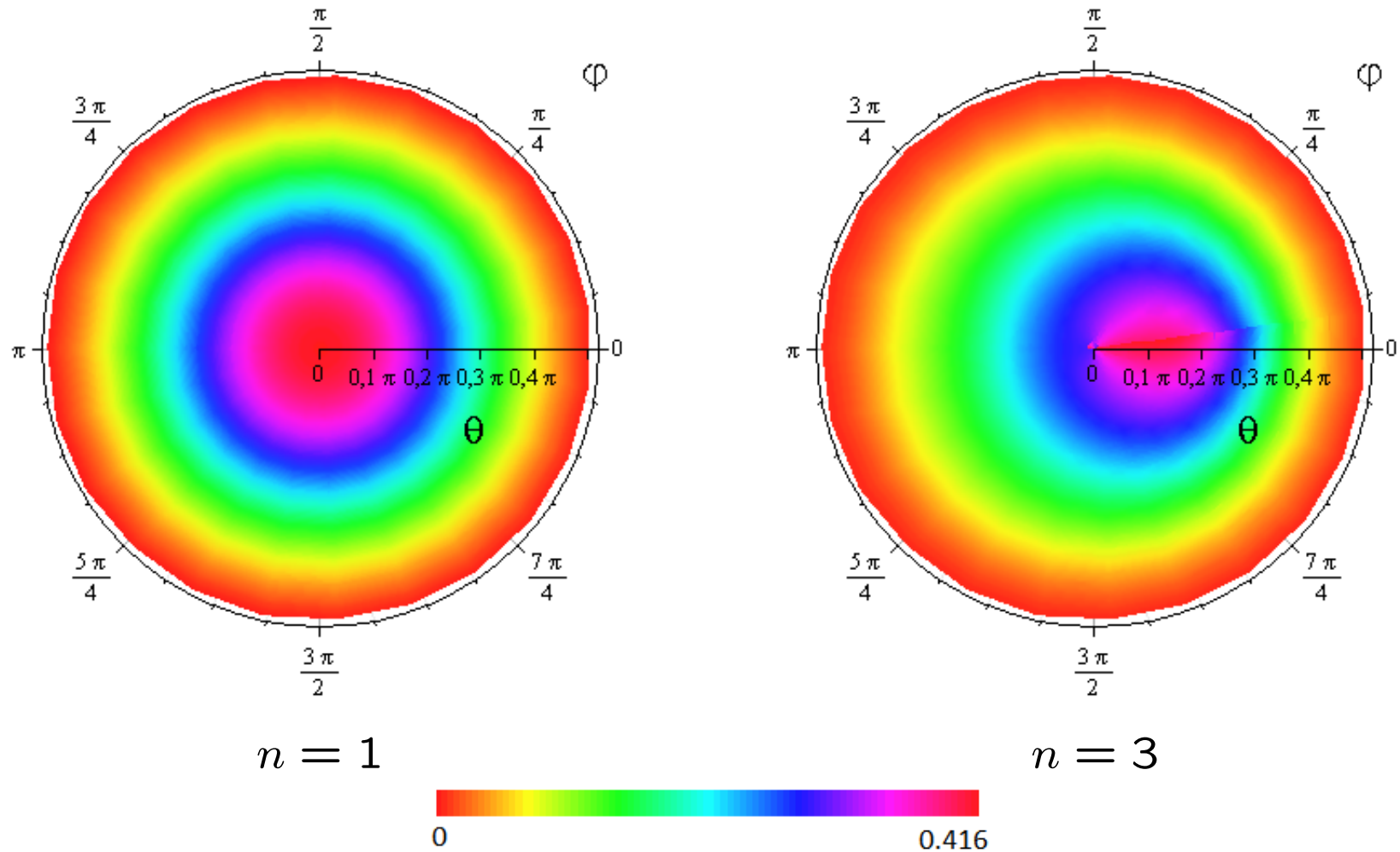


Angular distribution function of the normalized flux of radiation $\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$, backscattered by a layer of optical thickness $H = 5$. Aerosol scattering. $\theta_0 = 45^\circ$, $\varphi_0 = 0$. Hemispherical basis.

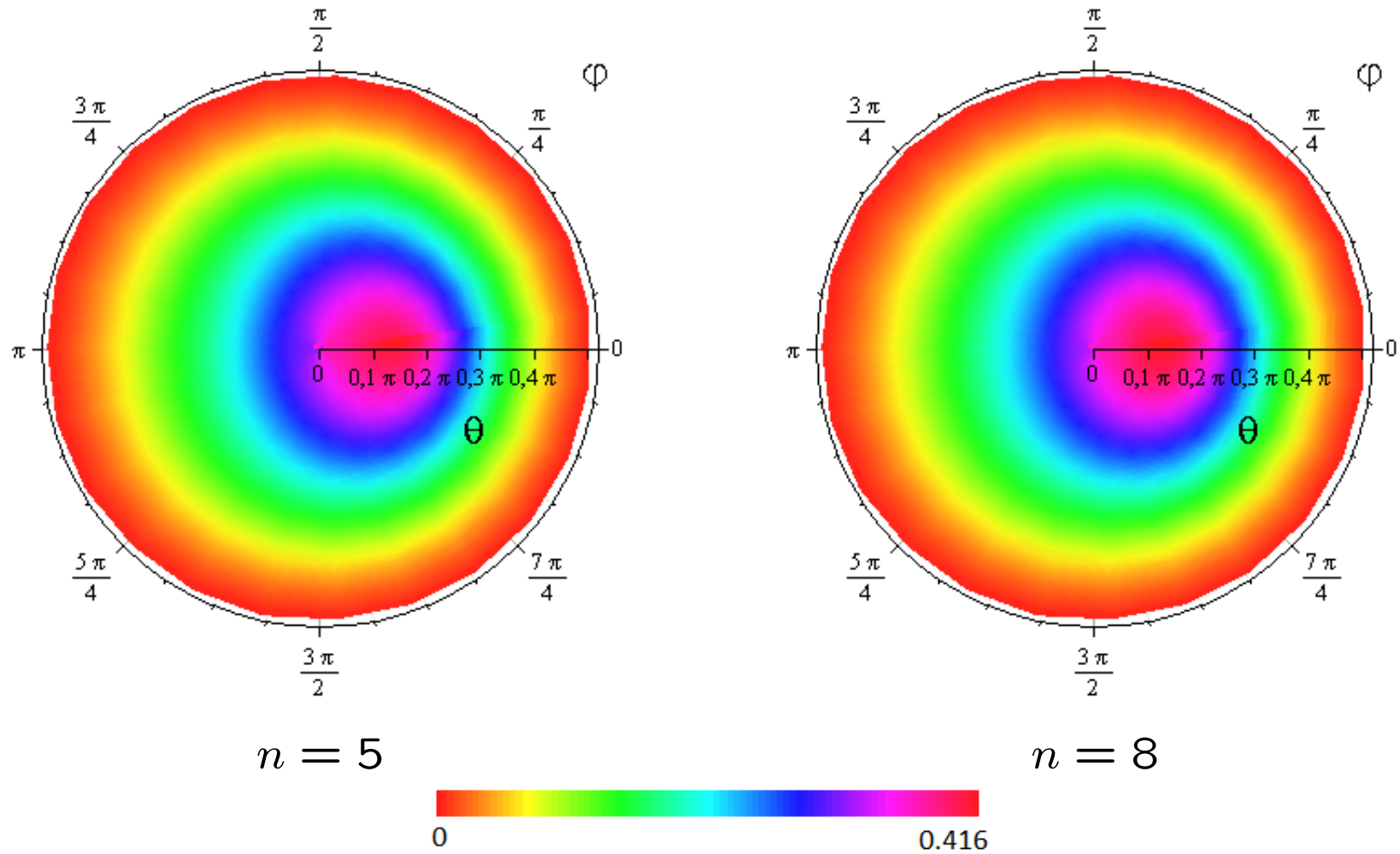
28



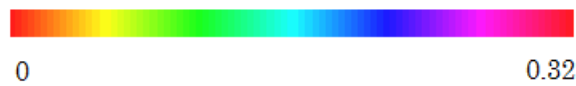
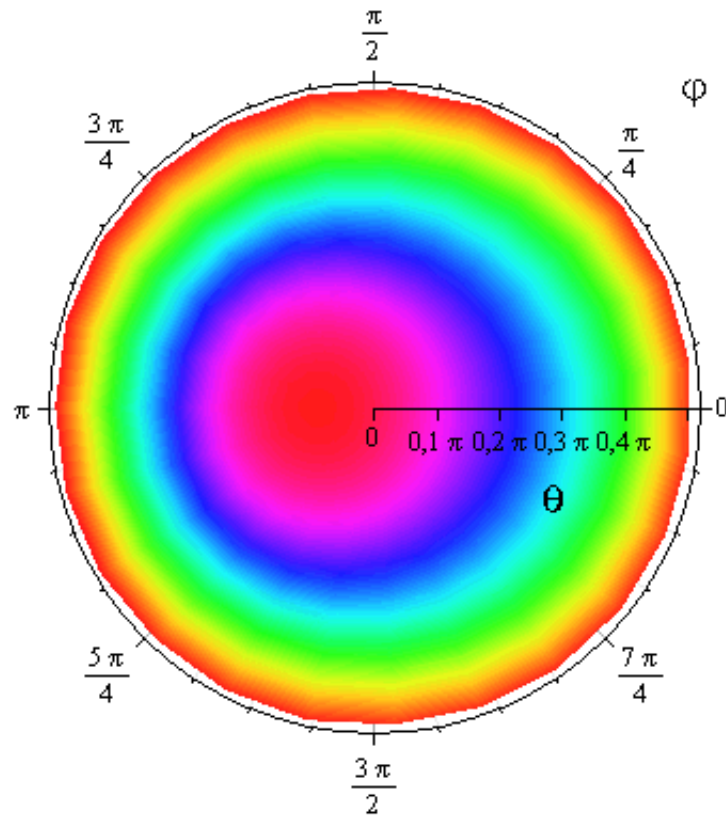
Angular distribution function of the normalized flux of radiation $\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$, transmitted by a layer of optical thickness $H = 5$. Aerosol scattering. $\theta_0 = 45^\circ$, $\varphi_0 = 0$. Factorized basis, $n_i = n_j = n$. 29



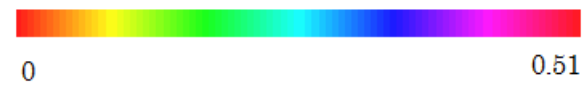
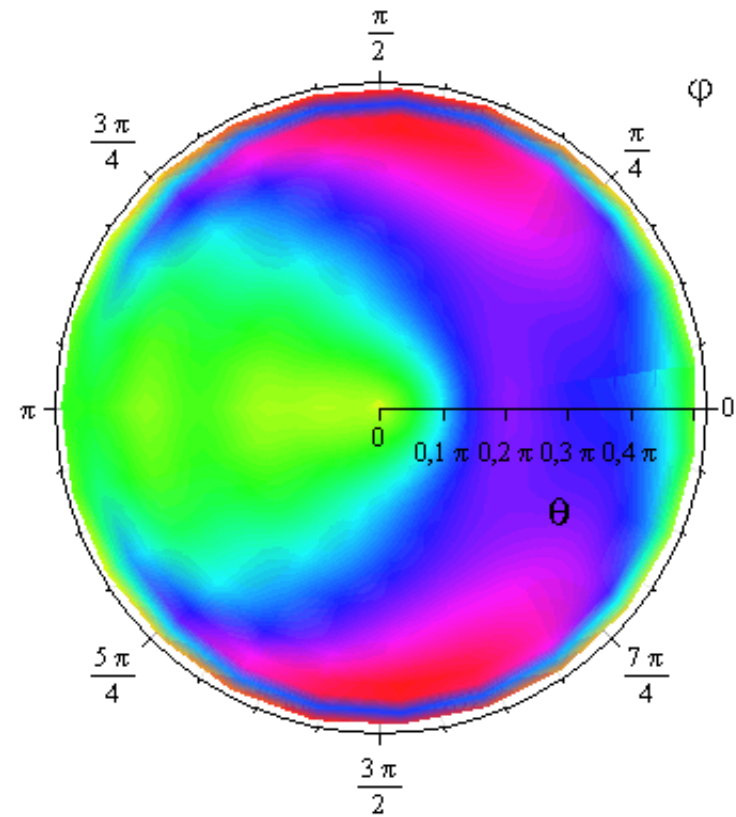
Angular distribution function of the normalized flux of radiation $\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$, transmitted by a layer of optical thickness $H = 5$. Aerosol scattering. $\theta_0 = 45^\circ$, $\varphi_0 = 0$. Factorized basis, $n_i = n_j = n$. 30



Angular distribution functions of the normalized flux of radiation $\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$ and the degree of polarization $p(\cos(\theta), \varphi)$, backscattered (1-st row) and transmitted (2-d row) by a layer of optical thickness $H = 5$. Molecular scattering. $\theta_0 = 45^\circ$, $\varphi_0 = 0$. Hemispherical basis, $n = 5$. 31

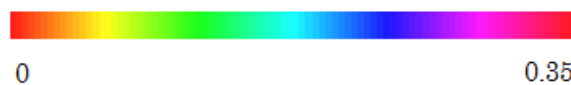
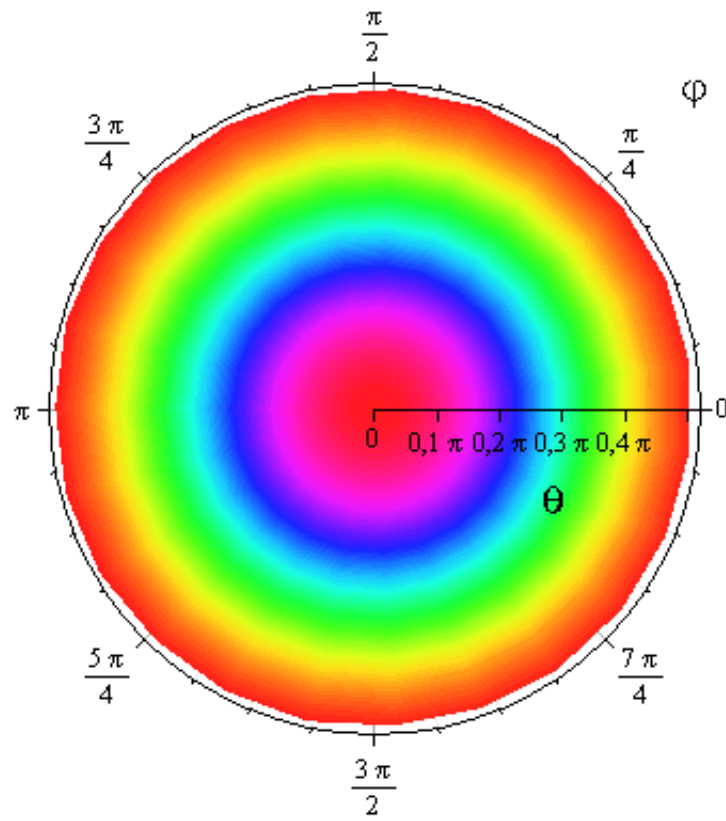


$$\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$$

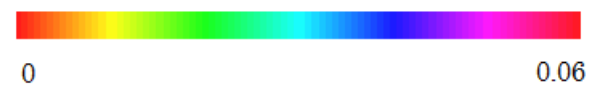
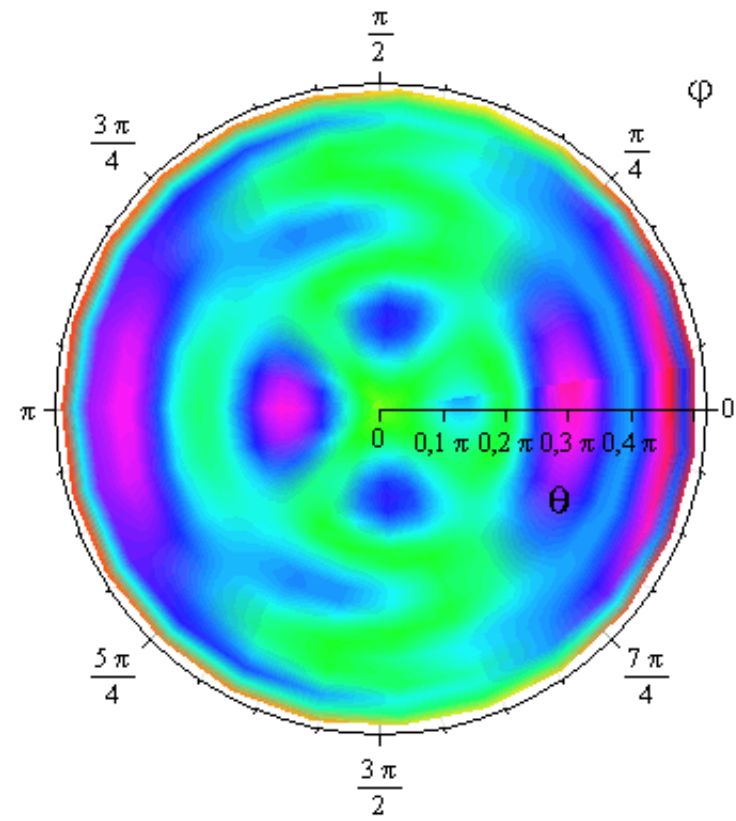


$$p(\cos(\theta), \varphi)$$

Angular distribution functions of the normalized flux of radiation $\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$ and the degree of polarization $p(\cos(\theta), \varphi)$, backscattered (1-st row) and transmitted (2-d row) by a layer of optical thickness $H = 5$. Molecular scattering. $\theta_0 = 45^\circ$, $\varphi_0 = 0$. Hemispherical basis, $n = 5$. 32

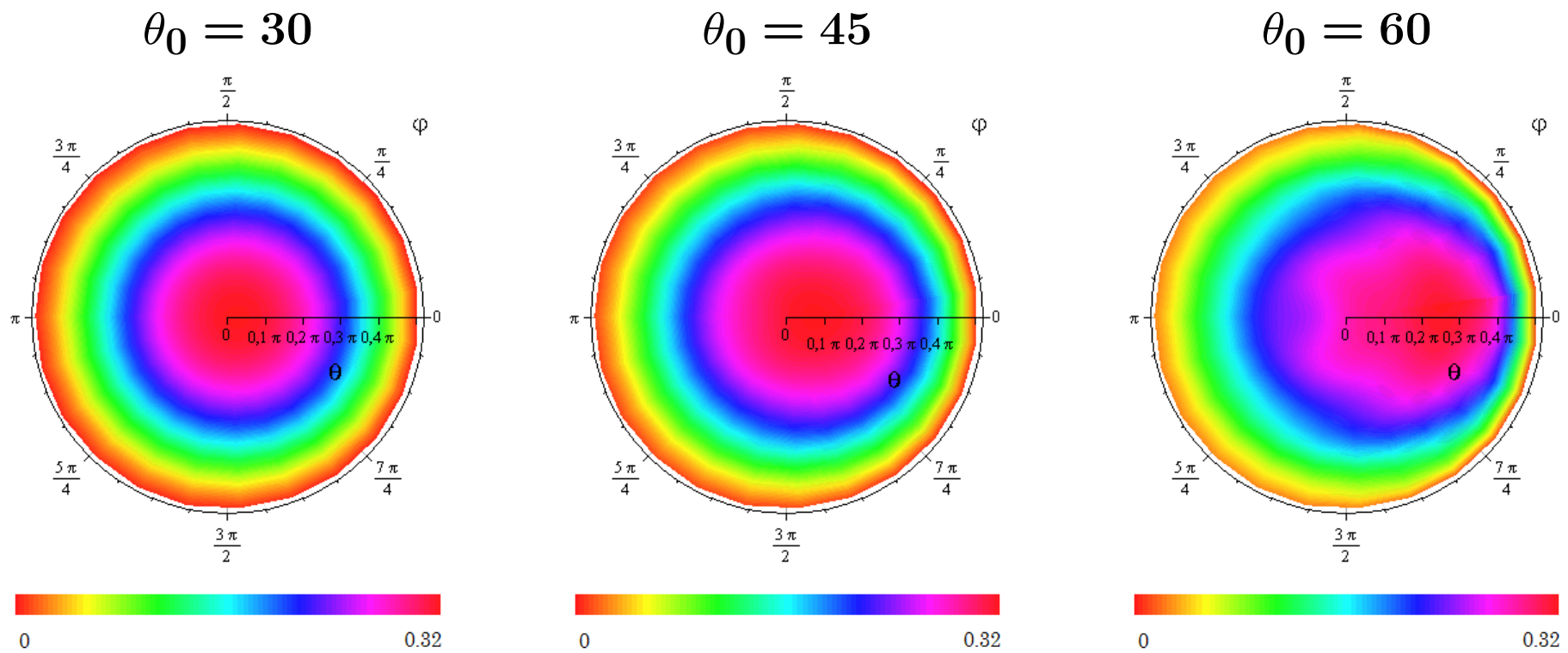


$$\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$$

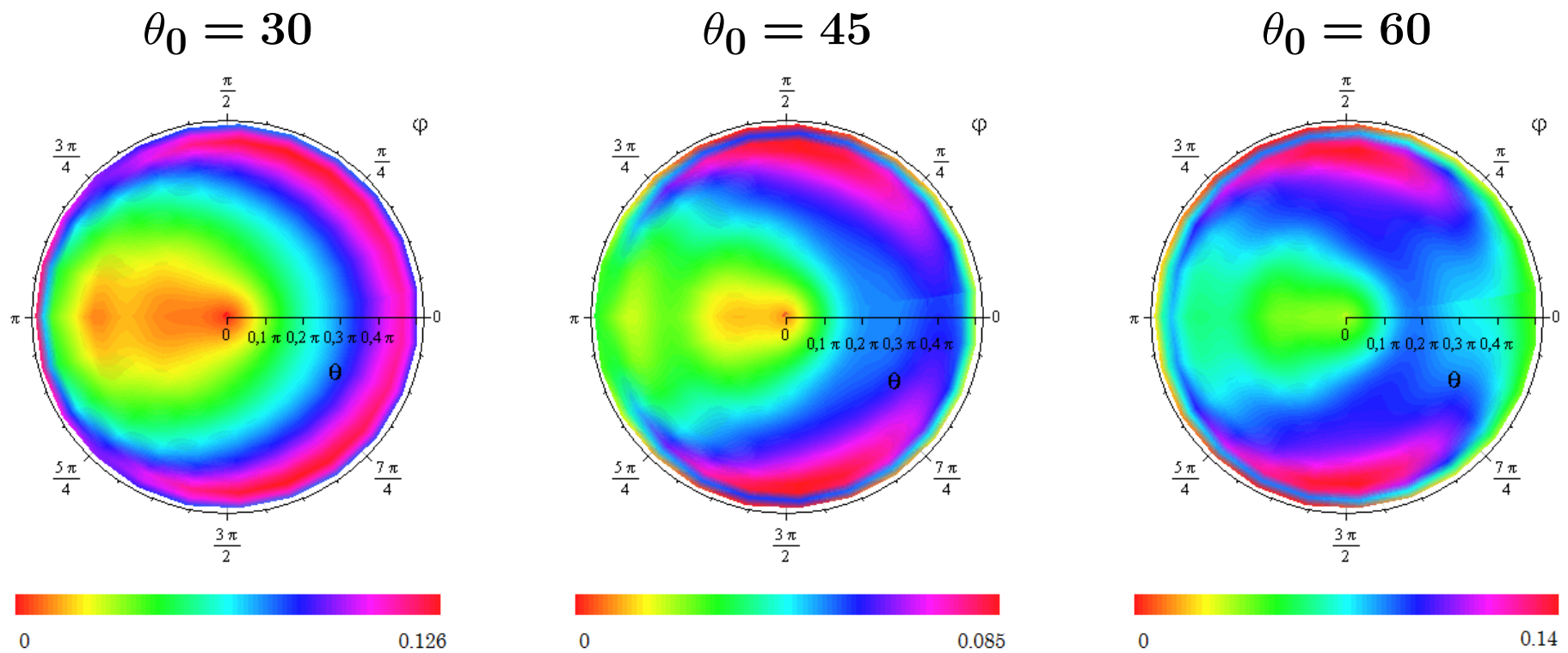


$$p(\cos(\theta), \varphi)$$

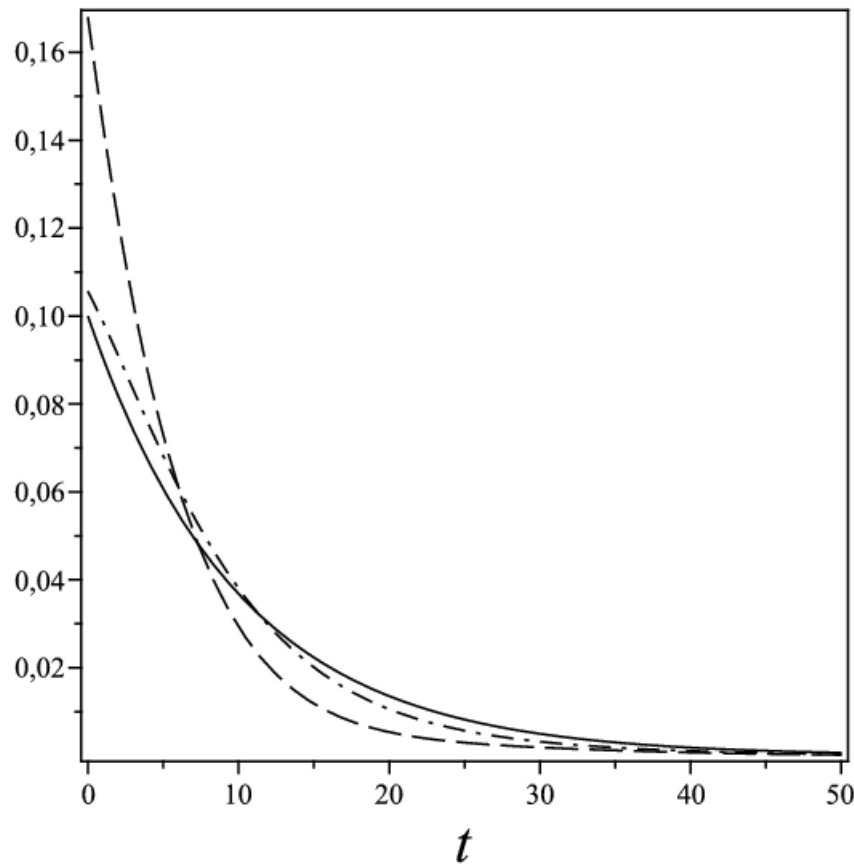
Angular characteristics of the radiation, backscattered by the layer of optical thickness $H = 10$. (a) – (c) The normalized radiation flux $\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$. (d) – (f) The degree of polarization $p_n(\cos(\theta), \varphi)$. $\varphi_0 = 0$. Hemispherical basis.



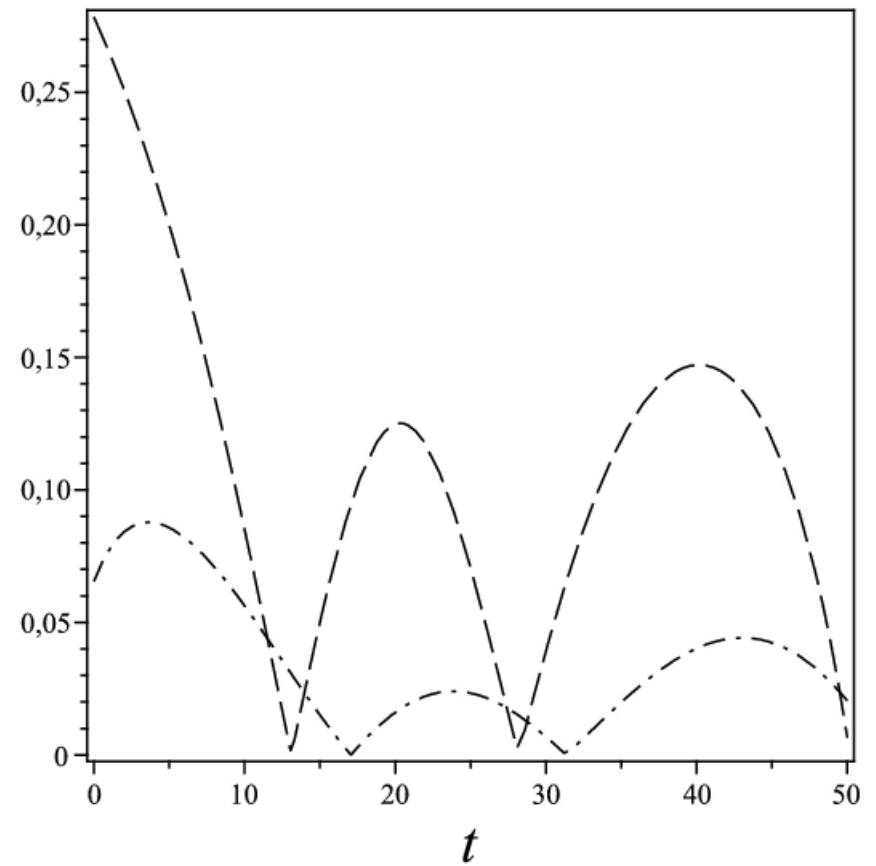
Angular characteristics of the radiation, backscattered by the layer of optical thickness $H = 10$. (a) – (c) The normalized radiation flux $\Phi_s^{(1)}(\cos(\theta), \varphi)/P_h$. (d) – (f) The degree of polarization $p_n(\cos(\theta), \varphi)$. $\varphi_0 = 0$. Hemispherical basis.



$\sigma_c = 0.1$. Solid line is for an exponential curve, dashed line is for molecular scattering matrix curve, dash-and-dot line is for aerosol scattering matrix curve. (a) Time distribution of radiation intensity, normalized over total flux, in comparison to an exponential asymptotics $e^{-\sigma_c vt}$. (b) Degree of polarization.

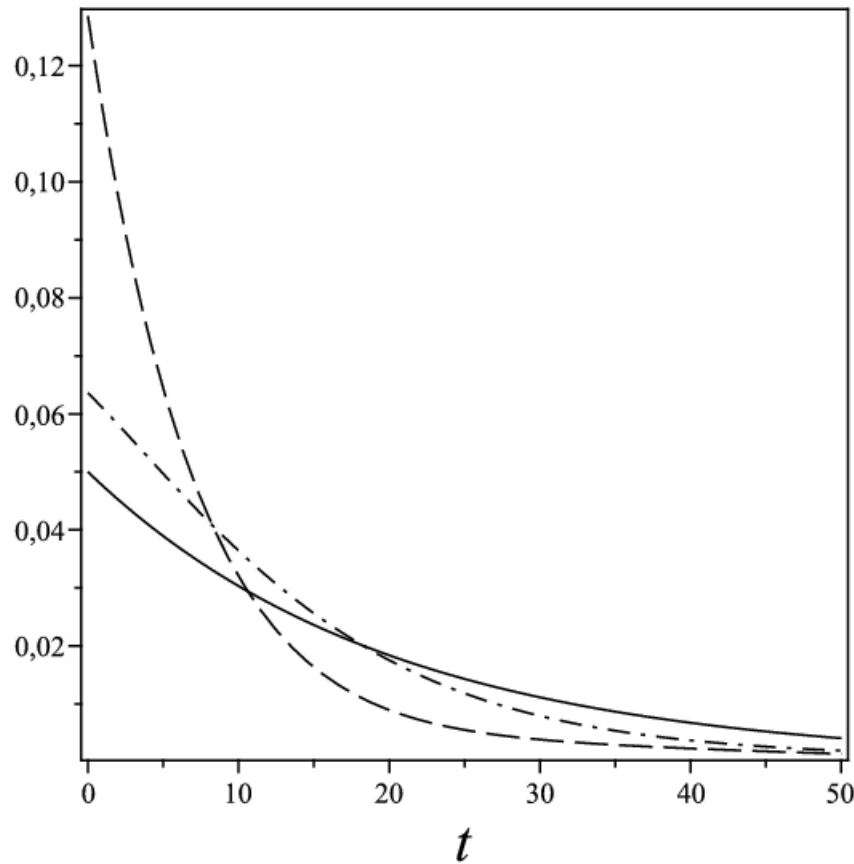


(a)

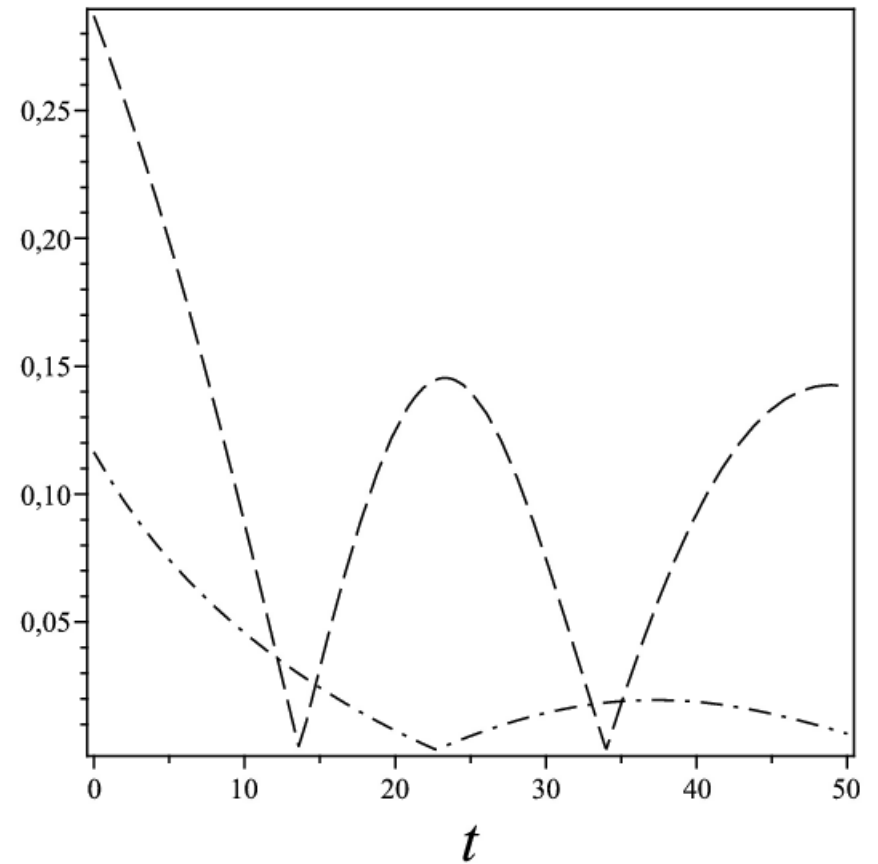


(b)

$\sigma_c = 0.05$. Solid line is for an exponential curve, dashed line is for molecular scattering matrix curve, dash-and-dot line is for aerosol scattering matrix curve. (a) Time distribution of radiation intensity, normalized over total flux, in comparison to an exponential asymptotics $e^{-\sigma_c vt}$. (b) Degree of polarization.



(a)



(b)

Thank you for attention!